# PHYSICS 241 Spring 1993 Prof. Bernice Durand COURSE INFORMATION

COURSE: Introduction to modern physics, lectures MWF 9:55, 3331 Sterling.

TEXT: Elementary Modern Physics, Paul A. Tipler, Worth Publishers 1992. There is a study guide which I didn't order at the bookstore. It looks good, but the text is already enough to read, in my opinion.

DISCUSSIONS: TUESDAYS, 1:20 or 2:25, 3331 Sterling, MANDATORY!

INSTRUCTOR: Prof. Bernice Durand, 4205 Chamberlin, 262-3827.

GOALS: We will master the textbook, which is a superb introduction to 20th century physics and very up-to-date. We will also experiment with different learning techniques and compare them, which means you must attend all lectures and one regular discussion section every week.

GRADE: There will be 42 lectures. In discussions and homework you will do work for credit averaging 2% of your grade per lecture. Thus 84% of your grade will be earned by steady, regular mastery of the material. A comprehensive exam on May 13 will count as 16% of your grade. Of the 84%, I estimate that 28% will be homework, much of it done in groups, 28% group work in discussions, and 28% individual work on quizzes given in discussions. I will instate hour exams only if the system doesn't seem to be working.

My usual expectations for final grades are A 90-100, AB 87-89, B 75-86, BC 72-74, C 60-71, D 50-59. I have no preconceptions as As h

D 50-59. I have no preconceptions as to how many of each grade I will give.

FORMAT: The 42 MWF lectures will cover the text approximately as shown on the schedule. Lectures will cover an average of 7 to 8 pages each. Some lectures will include trips to view demonstrations or apparatus, so you should be prompt.

There will be 14 Tuesday and one Friday (May 7) discussion sections, which are essential to your learning and to your grade. In these sessions, you will ask questions, do group problem-solving for credit and do individual problem-solving for credit, which will count as mini-exams. DO NOT MISS DISCUSSIONS!

Homework will be assigned on Wednesdays, in the first lecture on each chapter, and due on Wednesdays in the first lecture on the next chapter. Instructions for homework will include whether you are to work it in groups or alone. A grader will grade the homework for credit, with all members of a group receiving the same grade.

A comprehensive final exam for credit will cover the whole course.

GETTING HELP: Discussion sections and your study/working groups are for getting help. I will certainly be available after classes, MWF at 10:45 and Tues. at 3:15. I prefer to have you make an appointment with me for the next mutually agreeable time, so I don't have to sit in my office for certain hours with nobody showing up.

### PHYSICS 241, Spring 1993, Prof. Bernice Durand COURSE SCHEDULE

<b>WEEK</b> 1	MONDAY No class	TUESDAY Jan 19 startup	WEDNESDAY Jan 20 Secs 1.1-1.3	FRIDAY Jan 22 Sec 1.4
2	Jan 25	Jan 26	Jan 27	Jan 29
	Sec 1.5	Secs 1.1-1.5	Secs 1.6-1.8	Secs 1.9-1.10
3	Feb 1	Feb 2	Feb 3	Feb 5
	Sec 1.11	Secs 1.6-1.11	Secs 2.1-2.2	Secs 2.3-2.5
4	Feb 8	Feb 9	Feb 10	Feb 12
	Secs 2.5-2.6	Secs 2.1-2.6	Secs 3.1-3.2, App B	Secs 3.3-3.6
5	Feb 15	Feb 16	Feb 17	Feb 19
	Sec 3.7	Secs 3.1-3.7, App B	Secs 3.8-3.9	Sec 3.10
6	Feb 22	Feb 23	Feb 24	Feb 26
	Sec 3.11	Secs 3.8-3.11	Secs 4,1-4,2	Sec 4.2
7	Mar 1	Mar 2	Мат 3	Mar 5
S	Secs 4.3-4.4	Secs 4.1-4.4	Sec 4.5	Sec 4.6
Spring Recess	X	X	X	X
	Mar 15	Mar 16	Mar 17	Mar 19
	Secs 4.7-4.9	Secs 4.5-4.9	Sec 5.1	Secs 5.2-5.3
9	Mar 22	Mar 23	Mar 24	Mar 26
	Sec 5.3	Secs 5.1-5.3	Secs 6.1-6.2	Sec 6.2, App C
10 :	Mar 29	Mar 30	Mar 31	Apr 2
	Sec 6.3	Secs 6.1-6.3, App C	Secs 6.4-6.5	Secs 6.6-6.7
.11	Apr 5 Sec 6.8	Apr 6 Secs 6.4-6.8	Apr 7 Secs 7.1-7.2	No class
12	Apr 12	Apr 13	Apr 14	Apr 16
	Secs 7.3-7.4	Secs 7.1-7.4	Sec 7.5	Sec 7.6
13	Apr 19	Apr 20	Apr 21	Apr 23
	Secs 7.5-7.6	Secs 7.5-7.6	Secs 8.1-8.3	Secs 8.4-8.5
14	Apr 26	Apr 27	Apr 28	Apr 30
	Secs 8.6-8.8	Secs 8.1-8.8	Secs 9.1-9.2	Secs 9.3-9.5
15	May 3	May 4	May 5	May 7
	Secs 9.6-9.7	Secs 9.1-9.7	Sec 9.8	Discuss Sec 9.8
16	x	X	May 12 Review Session	May 13 Thursday Exam 2:45

# **Contents**

(Sections marked \* are optional.)

of F. Blatt. Physics-

Preface xvii

#### Chapter 1

# The Theory of Special Relativity I: The Lorentz Transformation 1

- 1.1 Introduction 1
- 1.2 Classical Relativity: The Galilean Transformation Equations 2
- 1.3 Electromagnetic Waves and the Luminiferous Ether 4
- 1.4 The Michelson-Morley Experiment 5
- 1.5 The Theory of Special Relativity 10
- 1.6 The Lorentz Transformations 12
  - 1.6(a) Simultaneity, Length Contraction, and Time Dilation 14
  - 1.6(b) The Twin Paradox 19
  - 1.6(c) The Velocity Transformations 21
- 1.7 Consequences of the Lorentz Transformations 22
  - 1.7(a) The Relativistic Doppler Effect 22
  - 1.7(b) Experimental Evidence of Relativistic Kinematics 24
- 1.8 The Relativistic Expressions in the Classical Limit 25

SUMMARY 26 BIBLIOGRAPHY 27 PROBLEMS 27

Chapter 2

# The Theory of Special Relativity: Relativistic Dynamics 30

- 2.1 Introduction 30
- 2.2 Relativistic Momentum 31
- 2,3 Energy 34
- 2.4 Relativistic Invariants 38
- 2.5 Force and Acceleration 42

SUMMARY 45 BIBLIOGRAPHY 45 PROBLEMS 46

Chapter 3 *The General Theory of Relativity	48
3.1 Introduction 48	
3.2 The Principle of Equivalence 49	

and Length Contraction 52
3.4 The General Theory of Relativity: Gravitation 56

3.3 Gravitational Red Shift: Gravitational Time Dilation

3.5 Predictions of the General Theory of Relativity 56

SUMMARY 62 BIBLIOGRAPHY 62 PROBLEMS 63

#### Chapter 4

#### Roots of the Quantum Theory 64

4.1 Introduction 64

4.2 Blackbody Radiation 654.2(a) Derivation of the Planck Distribution Law 67

4.3 Specific Heat 72 - 4.3(a) Specific Heat of Crystals 72 4.3(b) Specific Heat of Gases 76

4.4 The Photoelectric Effect 77

4.5 X Rays 80

4.6 Compton Scattering 84

SUMMARY 87 BIBLIOGRAPHY 88 PROBLEMS 88

#### Chapter 5

#### The Bohr-Rutherford Nuclear Atom 90

5.1 Charge and Mass of an Electron 90

5.2 Scattering Cross Section 93

5.3 Coulomb (Rutherford) Scattering 95

5.4 The Bohr Model of the Hydrogen Atom 99

5.5 The Energy-Level Diagram; Emission and Absorption of Radiation 103

5.6 Characteristic X-Ray Lines 106

5.7 The Franck-Hertz Experiment 109

5.8 The Correspondence Principle 111

SUMMARY 112 BIBLIOGRAPHY 113 PROBLEMS 113

#### 'Chapter 6

#### The Wave Nature of Particles 115

6.1 Introduction: The de Broglie Relation 115

6.2 Experimental Evidence of Electron Waves 116

6.3 Complementarity III	119	Complementarity	6.3
-------------------------	-----	-----------------	-----

6.4 The Uncertainty Principle 119

6.5 The Wave-Particle Duality and Complementarity: A Gedanken Experiment 125

SUMMARY 128 BIBLIOGRAPHY 129 PROBLEMS 129

#### Chapter 7

#### The Schrödinger Equation 131

- 7.1 Introduction 131
- 7,2 The One-Dimensional Schrödinger Equation 132
- 7.3 The Time-Independent Schödinger Equation 134
- 7.4 Interpretation of the Wave Function: Probability Density and Expectation Values 135
- 7.5 Wave Packets: Group and Phase Velocities 137
- 7.6 Particle in a One-Dimensional Square Well 140
   7.6(a) Infinite Potential Barriers 140
   7.6(b) Finite Potential Barriers 143
- 7.7 Parity 147
- 7.8 Tunneling 148
- 7.9 The Harmonic Oscillator 149

SUMMARY 155 BIBLIOGRAPHY 156 PROBLEMS 156

#### Chapter 8

### The Schrödinger Equation in Three Dimensions: The Hydrogen Atom 159

- 8.1 Introduction 159
- 8.2 Solution of the Schrödinger Equation in Spherical Coordinates 160
  - 8.2(a) Probability Densities and Expectation Values 162
- 8.3 Angular Momentum in Quantum Mechanics 1678.3(a) Spatial Quantization 169
- 8.4 Degeneracy 171

SUMMARY 172 BIBLIOGRAPHY 173 PROBLEMS 173

#### Chapter 9

#### Intrinsic Spin of the Electron 174

- 9.1 The Zeeman Effect and Electron Spin 174
  - 9.1(a) The Zeeman Effect 174
  - 9.1(b) Electron Spin 178

9.2	The Pauli Exclusion Principle and the Structure
	of Many-Electron Atoms 179
	9.2(a) The Ground State of Atoms 180
	9.2(b) Relation between Ground-State Configuration and Chemical
	and Physical Properties of Atoms 182
9.3	Fine Structure, Spin-Orbit Coupling, and the Anomalous
	Zeeman Effect 185
9.4	The Exclusion Principle and the Symmetry Properties
	of Electronic Wave Functions 188
9.5	Masers and Lasers 191
	9.5(a) Spontaneous and Stimulated Emission 191

9.5(b) Creating a Population Inversion: Pumping 194

SUMMARY 199 BIBLIOGRAPHY 201 PROBLEMS 201

#### Chapter 10

#### Molecular Structure and Molecular Spectra 203

10.1 Introduction 203 10.2 Ignic Bonding 294

10.3 Covalent (Homopolar) Bonding 207 10.3(a) The H<sub>2</sub> Ion 208 10.3(b) The H<sub>2</sub> Molecule 210

9.5(c) Maser and Laser Operation 195

10.4 Molecular Energy Levels and Molecular Spectra 212
 10.4(a) Rotational Energy Levels 212
 10.4(b) Vibrational Energy Levels 215
 10.4(c) Molecular Spectra 216

SI/MMARY 221 BIBLIOGRAPHY 222 PROBLEMS 223

### $\mathcal{U}_{\mathfrak{O}}$

#### Chapter 11

#### Statistical Physics 224

11.1 Introduction 224

11.2 Classical Statistical Physics 224

11.3 Quantum Statistical Physics 23011.3(a) Bose-Einstein and Fermi-Dirac Distributions 230

11.4 Liquid Helium and Bose-Einstein Condensation 235 11.4(a) Bose-Einstein Condensation 240

SUMMARY 243 BIBLIOGRAPHY 244 PROBLEMS 244

Chapter 12	
Solid-State Physics I: Structure of Crystalline Solids	
and Electron Energy Bands 246	
./12.1 Introduction 246	
2.2 Cohesive Forces in Crystals 247	
12.2(a) Ionic Crystals 250	
12.2(b) Covalent Crystals 252	
12.2(c) Metallic Crystals 253	
12.2(d) Molecular Crystals 254	
12.3 Crystal Defects 255	z
*12.3(a) Dislocations; Plastic Deformation and Crystal Growth 25	D
12.3(b) Point Defects 258	
Impurity Atoms 258	
Vacancies and Interstitials; Solid-State Diffusion 259	
12.4 The Band Theory 260	
12.4(a) Bloch Functions 261	
12.4(a) Block Policions 257	
√12.5 Metals, Insulators, and Semiconductors 264	
*12.A Appendix: The Bloch-Floquet Theorem 268	
*12.B Appendix: Solution of the One-Dimensional Schrödinger	
Equation in the Tight-Binding Approximation 269	
SUMMARY 271	
BIBLIOGRAPHY 272	
PROBLEMS 272	
Chapter 13	
Solid-State Physics II: Electronic Properties 274	
/13.1 Introduction 274	
13.2 Metals 274	
13.2(a) The Electronic Specific Heat 274	
13.2(b) Electrical Conductivity 277	
13.2(c) Thermal Conductivity; the Wiedemann-Franz Law 281	
/13.3 Semiconductors 281	
13.3(a) Conductivity 284	
13.3(b) The Hall Effect 285	
13.4 Semiconductor Devices 287	
13.4(a) p-n Junction Diodes 287	
13.4(b) Tunnel Diodes 290	
13.4(c) Transistors 291	
The n-p-n Transistor 291	
Metal-Oxide-Semiconductor Field-Effect Transistor	
(MOSFET) 292	
13.4(d) Optical Devices 294	
Photovoltaic Effect 296	
Light-Emitting Diodes (LEDs) and Junction Lasers 297	
Tight-channik Diones (DDDs) and supplied theses and	

13.5 Superconductivity 298
13.5(a) Fundamental Properties of Superconductors 298
13.5(b) Superconductor Technology 303
Josephson Junctions 305
SQUID (Superconducting Quantum Interference
Devices) 306
Josephson Junction Logic Elements 309

SUMMARY 310 BIBLIOGRAPHY 312 PROBLEMS 312

### Chapter 14

#### Nuclear Physics 1: Properties of Nuclei 315

- 14.1 Introduction 315
- 14.2 Nucleons and Nuclei 318
  14.2(a) Composition of Nuclei: Protons and Neutrons 318
  14.2(b) Size and Shape of Nuclei 320
- 14.3 Nuclear Force 324 14.3(a) Origin of the Nuclear Force; the Exchange Force 326
- 14.4 Stability of Nuclei: Binding Energy 327
- 14.5 The Nuclear Shell Model 333
- \*14.6 Nuclear Magnetic Resonance (NMR) 335
  - 14.6(a) Principles of NMR 335
  - 14.6(b) Energy Absorption and Spin-Lattice Relaxation 337
  - 14.6(e) Nuclear Magnetization and the Dynamics of the Magnetization Vector 339
  - 14.6(d) Techniques for Observing NMR and Measuring  $T_1$  and  $T_2$  342
  - 14.6(e) Some Applications of NMR 346

SUMMARY 348 BIBLIOGRAPHY 349 PROBLEMS 350

#### Chapter 15

## Nuclear Physics II: Radioactivity and Nuclear Reactions 351

- 15.1 Introduction 351
- 15.2 Decay Constant, Half-Life, and Activity 352
- 15.3 Radioactive Decay Processes 361
  - 15.3(a) Alpha Decay 361
  - 15.3(b) Beta Decay 363
  - 15.3(c) Gamma Rays and Nuclear Energy Levels 366
- 15.4 Mössbauer Effect 367
- 15.5 Nuclear Reactions 370
  - 15.5(a) Reaction Cross Section 372
  - 15.5(b) Nuclear Spectroscopy 373
  - 15.5(c) Compound Nucleus 376

15.6 Fission and Fusion 378
15.6(a) Fission; Reactors 378
15.6(b) Fusion 385
\*15.7 Applications of Nuclear Physics 389
15.7(a) Radioisotopes in Medicine 389
Diagnosis 389
Radioisotopes for Therapy 394
15.7(b) Radioisotopes in Archeology 396

SUMMARY 402

SUMMARY 402 BIBLIOGRAPHY 404 PROBLEMS 404

\*Chapter 16

Instrumentation for Nuclear Research 407

16.1 Rarticle Accelerators 407
16.Na) Linear Accelerators 407
16.1(b) Cyclotrons and Synchrotrons 412

16,2 Particle Detectors 420

SUMMARY 426 BIBLIOGRAPHY 426 PROBLEMS 427

Chapter 17

#### Elementary Particles 428

17.1 Introduction 428

17.2 Relativistic Quantum Mechanics 429

17.3 Classification of Elementary Particles 431

17.4 Conservation Laws, Symmetry, and Selection Rules 434

17.5 Resonance Particles 438

17.6 More About Leptons 439

17.7 Quarks 441

17.8 Unified Field Theories 447

SUMMARY 448 BIBLIOGRAPHY 449 PROBLEMS 449

Appendix A :

The Fundamental Physical Constants 451

Appendix B

Table of Isotopes 453

Appendix C

Nobel Laureates in Physics 492

Answers to Odd-Numbered Problems 497

Index 501

#### PHYSICS 241 — SPRING 1994 — SYLLABUS

Professor Robert March

TEXT — Elementary Modern Physics, by Paul A. Tipler, Worth 1992.

Some additional readings, labeled by chapter, are included in this packet, and more may be handed out as the semester progresses.

The schedule of lectures below may be approximate, as this is the first time I have used this text in this course. Exams may cover material in lectures up to one week preceding the exam.

		SCHEDULE:	·
Week of:	Lectures:	·	Discussion:
Jan 24	Chapter 1 MWF	·	get-acquainted
Jan 31	Chapter 1 MWF		Problem Set 1 due
Feb 7	Chapter 2 MWF	· · · · · · · · · · · · · · · · · · ·	in-class problems A
Feb 14	Chapter 2 M,	Chapter 3 WF	Problem Set 2 due
Feb 21	Chapter 3 MWF		in-class problems B
Feb 28	Chapter 3 M,	Chapter 4 WF	6-week exam
Mar 7	Chapter 4 MWF		in-class problems C
Mar 14	Chapter 4 M,	Chapter 5 WF	Problem Set 3 due
Mar 21.	Chapter 5 M,	Chapter 6 WF	in-class problems D
Mar 28	SPRING	BREAK	
Apr 4	Chapter 6 WF		in-class problems E
Apr 11	Chapter 6 MWF		Problem Set 4 due
Apr 18	Chapter 7 MWF		12-week exam
Apr 25	Chapter 7 M,	Chapter 8 WF	in-class problems F
May 2	Chapter 8 MW,	Chapter 9 F	Problem Set 5 due
May 9	Chapter 9 MWF		Essay due, Review for final

Final Exam Wednesday, May 18 at 12:25 pm

#### OFFICE HOURS AND EMAIL

Office hours: 10-11:45am Wed., 1:30-3pm Thu., or call for appointment.

My office is 4289 Chamberlin, the first office past the bridge from the third floor of Sterling Hall to Chamberlin. My campus phone is 262-5947. My mailbox is near the University Avenue entrance to Chamberlin.

My email (Internet) address is bobmarch@wishep.physics.wisc.edu. This is a good way to get in touch with me, and I will try to respond rapidly to questions.

#### DISCUSSIONS, EXAMS, AND PROBLEMS

It is important to attend discussion sections. In nearly every week, something will be going on that could affect your grade. In all weeks, the first few minutes of discussion will be devoted to questions on lectures and reading.

The 6-week and 12-week exams will be given during the discussion hour. A study guide listing the topic of each question, with suggestions for study, including problems similar to those on the exam, will be handed out a week prior to each exam. NOTE – you may bring two sheets (written on both sides) of notes to each hour exam, four sheets to the final.

In weeks when no problem set is due, you will divide up into groups of three or four to work problems assigned on the spot. A single copy of your solutions will be turned in, and all members of the group will receive the same grade. Membership in groups will rotate. The purpose is to teach you to study together, a skill that will serve you well in more advanced courses. In weeks when a problem set is due, we will discuss the problems. You may be called upon to work a problem at the blackboard.

#### GRADING

Your grade will be based 70% on exams, 30% on problem sets.

The final exam will have two parts, one cumulative and the other on the last part of the course. Each will have a separate score, making for four exam scores in all. In computing your grade, these will be *performance-weighted*: your best score will count for 30%, second-best 20%, next 12%, lowest 8%.

Individual problem sets will count for 20% of your grade, and group problems worked in class for 10%. In each case, your worst score will be discarded. The essay (see next page) will be lumped with problem sets for grading purposes.

There is no "curve" in this course — i.e. I have no preconceptions as to how many of each grade I should give. My target for grade boundaries is generally A=92, AB=89, B=80, BC=75, C=65, and D=60. I may modify these a bit, but never upwards — you are NOT in competition with one another!

#### SHORT ESSAY

At the end of this packet you will find an article by Sylvan Schweber that I believe is a perceptive analysis of the central thrust of 20th century physics, and how that thrust is likely to be modified in the 21st century. It should be read in conjunction with Chapter 8. You should write a short (~ 3-page) essay addressing the following points:

- 1. Did you find the article understandable and informative?
- 2. What (if anything) in it particularly affected your thinking?
- 3. Was there anything in it with which you disagreed?

Your grade on this essay will be lumped with that of the short Problem Set 1 to make a single grade. The grade will be less than perfect only if you fail to address any of the questions above.

#### PROBLEM SETS 1 & 2

Set 1	Due February 1	Chap 1, prob 2, 3, 12, 31, 49
Set 2	Due February 15	Chap 1, prob 33b, 39, 41, 76
		Below, prob $A, B$
<u> </u>	·	Chap 2, prob 5, 7b

The problems below are based on the material on pages 4 & 5 of this packet:

Problem A — A particle has total energy 1000 MeV and momentum 870 MeV/c. Find its mass (rest energy) and its dimensionless velocity  $\beta$ .

Problem B op An unstable neutral particle of unknown mass decays into two particles with mass, energy, and momentum  $(M_1, E_1, P_1)$  and  $(M_2, E_2, P_2)$ . The opening angle between the particles' paths is  $\theta$ . Show that the mass of the parent particle is given by

$$M = [M_1^2 + M_2^2 + 2E_1E_2 - 2P_1P_2\cos\theta]^{\frac{1}{2}}$$

#### LATE PAPER POLICY

Late problem sets turned in by Friday of the week assigned will get 75% credit. Anything later will get 50% credit.

#### SUPPLEMENTARY READINGS FOR CHAPTER 1

#### Dimensionless Notation $(\gamma, \beta, \eta)$

Many relativity texts employ a notation in which the coefficients in the Lorentz transformation are dimensionless ratios. The Lorentz factor  $\gamma = E/mc^2$  is one of these. The other two are:

$$\beta = \frac{v}{c} \qquad \qquad \eta = P/Mc$$

The following identities are useful:

$$\eta=eta\gamma \qquad \qquad \gamma^2=\eta^2+1 \qquad \qquad \gamma=[1-eta^2]^{-rac{1}{2}}$$

With this notation, the Lorentz transformation becomes:

$$x = \gamma x' + \eta c t'$$
  $y = y'$   $z = z'$   $ct = \eta x' + \gamma c t'$ 

#### "Natural" Units

These are sometimes referred to as "units in which c = 1," but this statement is not as mysterious as it seems. Its simply means that in relativistic kinematics we measure both momentum and mass in energy units, i.e. the symbol "M" stands for the rest energy  $Mc^2$ , and "P" stands for "Pc." Since particle physics and astrophysics has no distinct units for momentum or mass, an energy unit, the electron-volt and its multiples, is employed for all three, which are related by the identity

$$E^2 = P^2 + M^2$$

In these units, momentum is stated in "eV/c," which is equal to Pc in eV.

EXAMPLE — A  $\pi$ -meson has mass (rest energy) 140 MeV. Hits momentum is 200 MeV/c, then its total energy is

$$E = \sqrt{P^2 + M^2} = \sqrt{200^2 + 140^2} \ MeV = 244 \ MeV$$

(A more ambitious refinement of this notation, employed in quantum field theory, has  $\hbar = c = 1$ , but it would only confuse matters to introduce it at this time.)

#### Four-dimensional Invariants

The Lorentz transformations for x and t convert what is a pure interval of time or space in one Lorentz frame is a mixture of a time interval and a space interval in any other Lorentz frame. In 1908 Hermann Minkowski (who had been one of Einstein's teachers in Zurich) noted that if time is treated somewhat as if it were a fourth dimension of space, the interval between two events is the same in all Lorentz frames. I use the phrase "somewhat as if" to emphasize that this formalism does not eliminate even the formal distinction between space and time, much less the psychological one: it is simply a "bookkeeping device" that gives a frame-independent representation of reality that is both appropriate and useful in relativistic systems.

Formally, it means: (1) measure both space and time in the same units, i.e. replace t by ct; (2) define a modified four-dimensional pythagorean length,

$$\delta s^2 = \delta x^2 + \delta y^2 + \delta z^2 - c^2 \delta t^2$$

(that minus sign on the time term is the basis for my "somewhat as if" remark.) This interval is the same in all Lorentz frames. For example, in the case of the "lightning-struck train" in Section 1-5 of the text, subtracting the square of time interval  $L_pV/c^2$  from the square of the distance between the lightning strikes in the train rest frame gives us  $L_p^2$ , because in the platform rest frame the time interval is zero. Put another way, the length of the "four-vector"  $(\delta x, \delta y, \delta z, c\delta t)$  is invariant under the Lorentz transformation.

What makes this formalism useful is that there are many other such four vectors, one of which is the energy-momentum vector (usually called the four-momentum, which in natural units is written

$$\vec{Q} = (P_x, P_y, P_z, E)$$

To compute the magnitude of Q, we subtract the squares of the momentum components. For a single particle, this gives the rest mass. For a system of several particles that originates from the decay of an unseen unstable particle, this invariant is the rest mass of the parent particle,

$$M^2 = (\Sigma E)^2 - |\Sigma \vec{P}|^2$$

In Problem B on Problem Set 2 you will express this quantity for the case of a two-body decay.